

AI as a Tool in Computational Astrophysics

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The background of the slide is a stunning astronomical image. It features a vast field of stars against a deep blue cosmic background. In the foreground, there are large, billowing clouds of interstellar dust and gas, primarily in shades of orange, red, and yellow, which appear to be illuminated by nearby stars. The clouds have a complex, filamentary structure with various ridges and valleys. The overall effect is one of immense scale and beauty in the universe.

Astrophysics: A shared computational challenge

Credit: NASA JWST

The Data Challenge

Understanding the physical processes governing the universe using:

Observations & Simulations

- Simulations & Observations: Massive scale (TB-PB-scale outputs)
- Non-linear, statistically noisy data & extensive parameter spaces
- Traditional analysis methods overwhelmed
- **Need:** Automated intelligence processing

The Data Challenge → Why AI in Astro

Understanding the physical processes governing the universe using:

Observations, simulations, and theory

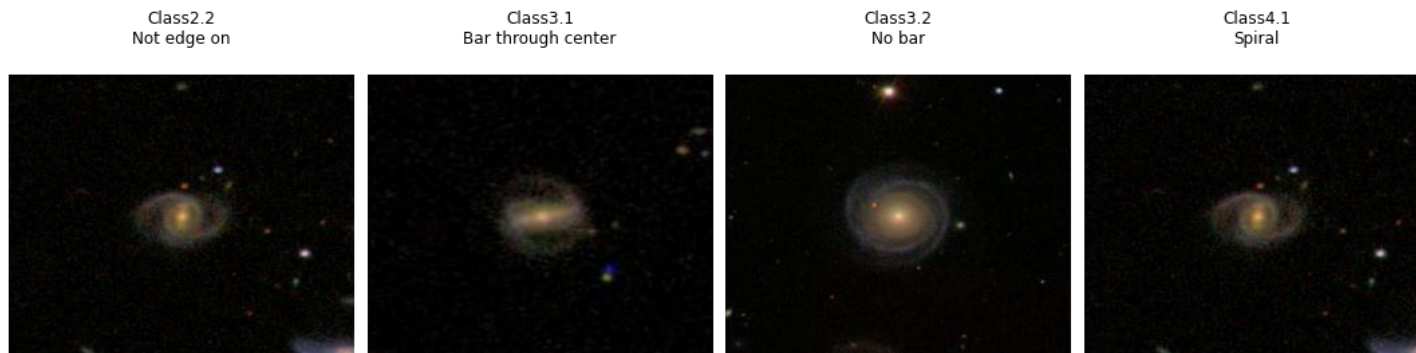
- Simulations & Observations: Massive scale (TB-PB-scale outputs)
- Highly correlated, high-dimensional data
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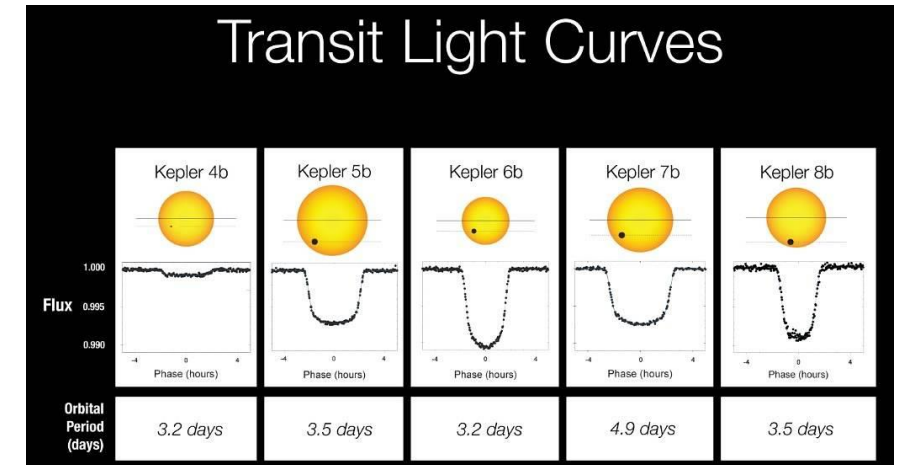
- Highly correlated, high-dimensional data
 - Pattern recognition
- Scalable to massive datasets
- Find non-obvious correlations
 - Self-/Un-supervised learning

Example AI applications in astrophysics

- Exoplanet detection in light curves
- Galaxy morphology classification
- Multi-modal searching – e.g. AstroCLIP
- Subgrid physics emulators



Jay Speidell, Galaxy Zoo



NASA/Kepler Mission



Traditional calculations vs. AI-enhanced calculations

Credit: JWST NIRCam image release (Chameleon I dark molecular cloud)

Strengths & weaknesses: Traditional algs

Strengths

- Physically motivated, interpretable
- Well-characterized uncertainties
- Extrapolates beyond training regimes
- Established validation/publication standards

Weaknesses

- $\uparrow$ complexity = $\uparrow\uparrow$ computational expense
- Programmed and constrained physics
- Misses unexpected non-obvious patterns
- Difficult to scale

Strengths & weaknesses: Modern AI algs

Strengths

- Efficiently handles high-dimensional feature spaces
- Discovers non-obvious correlations
- Scales efficiently
- Learns complex nonlinear relationships
- Rapid inference after training

Weaknesses

- Needs large training datasets
 - Labeled data can be expensive
- Interpretability is often difficult
 - Uncertainty quantification imprecise
- Often fails when extrapolating beyond training parameter space

Decision criteria: data volume, complexity,
interpretability, computational resources

Best approach: hybrid implementation

Physics-informed networks, ML for bottlenecks and traditional for
interpretation, Emulators validated against resolved simulations

Science First: Planning Projects

Decision criteria: data volume, complexity, interpretability, computational resources

Best approach: hybrid implementation

Physics-informed networks, ML for bottlenecks and traditional for interpretation, Emulators validated against resolved simulations

- Define clear science question first
 - Does your dataset provide a specific advantage?
- Determine where ML truly adds value
- Remain flexible on architecture choices

Science First: Planning Projects

Key Takeaways

- AI transforms astrophysics through scale
- Know the strengths and weaknesses: traditional & AI
- Hybrid approaches often work best
- Science question drives method choice
- Apply ML strategically to different data bottlenecks



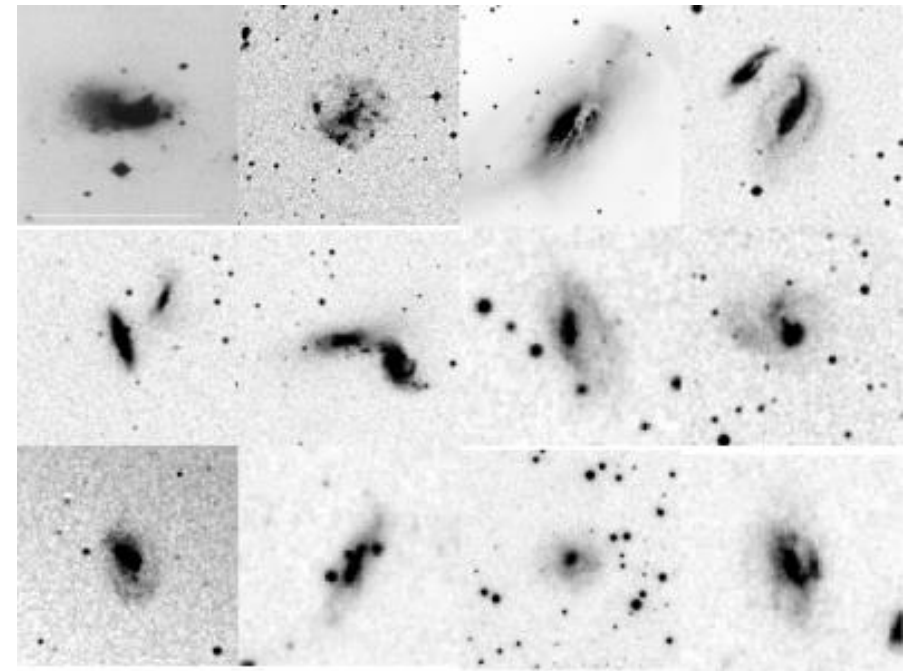
Credit: NASA JWST

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Thank you

Astrophysics overview

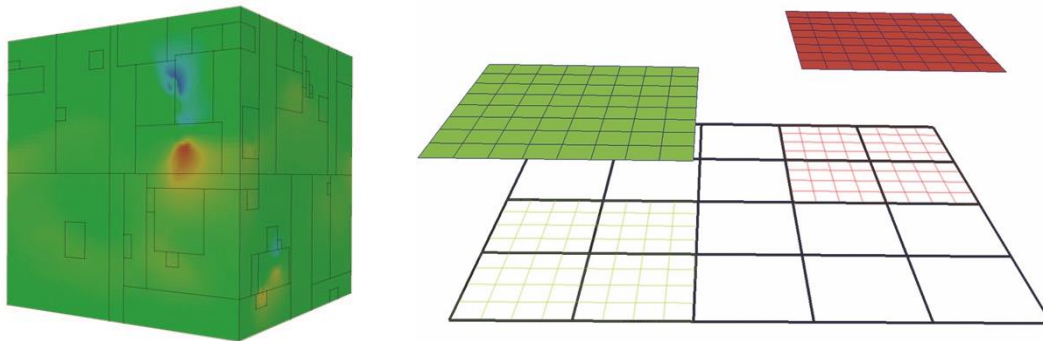
- Observation-specific challenges:
 - Instrumental noise and systematic errors
 - Incomplete sky coverage and selection effects
 - Limited time resolution
 - No control over astrophysical “experiments”



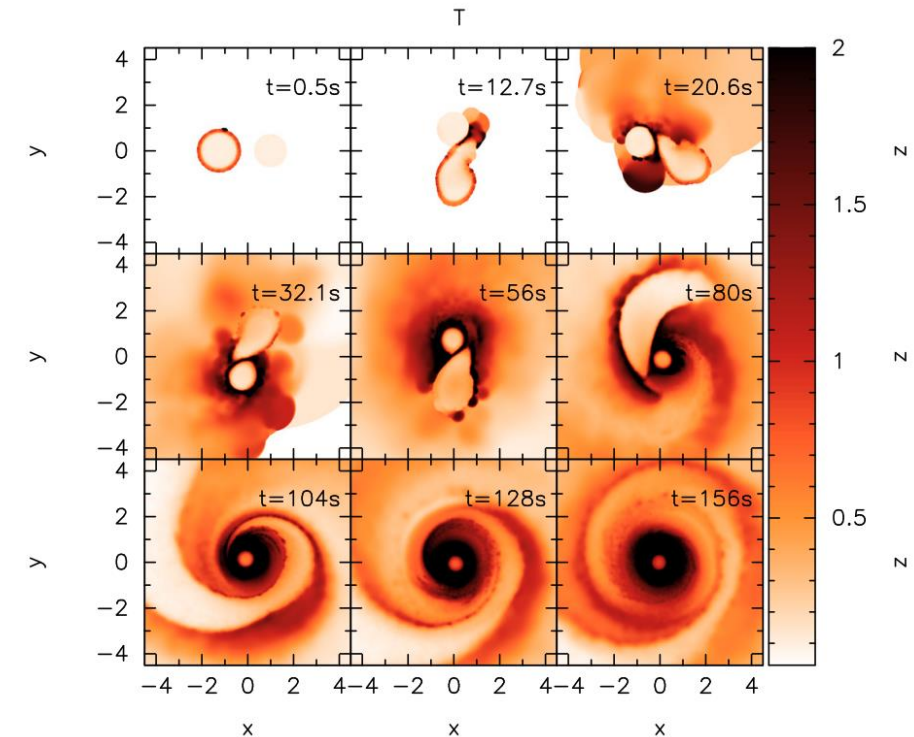
Conselice (2006)

Astrophysics overview

- Simulation-specific challenges:
 - Finite resolution limits resolved physics
 - Subgrid physics requires approximations
 - Computational cost limits parameter exploration



Sharkey et al. (2011)



Loren-Aguilar et al. (2009)

Data challenges made for AI

- **Scale and Volume**
 - PB+ scale datasets from surveys and simulations (LSST, Euclid, JWST; IllustrisTNG, FIRE, CAMELS)
 - Millions-billions of objects requiring classification (stars, galaxies, AGN, clusters, cosmic rays, binary systems, planetary systems, etc.)
 - Real-time processing of incoming data streams (telescope data; dynamic simulation modeling)
- **Complexity**
 - High-dimensional parameter spaces
 - Non-linear relationships between observables and physical properties
 - Complex analytical relationships in data (stellar history assumptions in observation, subgrid modeling in simulations)

Data challenges made for AI

- Computational bottlenecks
 - Computationally expensive to explore full parameter space + Multiphysics
 - Cross-cell flux calculations, subgrid approximations, implicit time-stepping schemes, self-gravity solvers replaceable with emulators “cheaply”
 - Creating mock observations (radiative transfer calculations like ray-tracing) requires thousands of iterations
- Pattern recognition
 - Identifying rare or unexpected phenomena in large datasets
 - Extracting weak signals from noisy observations
 - Morphological classification of complex structures

When to choose which tool

Decision criteria: data volume, complexity, interpretability, computational resources

Use traditional methods when:

- Studying simple well-understood processes
- Small datasets or rare phenomena
- Interpretability and transparency critical
- Extrapolation beyond data required
- Established methods work accurately and efficiently

Use AI methods when:

- Complex patterns in high dimensions
- Massive datasets requiring scalability
- No clear analytical model exists/is resolved
- Find unknown/missing correlations
- Large amounts training data available

Best approach: hybrid implementation

- Physics-informed networks with constraints
- ML for bottlenecks, traditional for interpretation
- Emulators validated against resolved simulations, real observations